# What Satellites See: Eyes Above the Skies 

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## How to measure the energies of very high energy nuclei.



Heavy nuclei were discovered by people at the U of Minnesota before I arrived.

They put what was basically a pile of photographic emulsions on a balloon flight in 1948 and saw interacting protons like the one at the left and heavy nuclei that made denser tracks.

Ionization loss $Z^{2}$

# A "star" from a nuclear 

 reaction - protons and other nuclei do interact with nuclei.The heaviest tracks are due to heavy nuclei, the medium density tracks are due to low energy protons ejected in the collision with an emulsion silver nucleus. The very light tracks are due to mesons (->
muons).

## Cover boy: the only cover I ever made

## THE MINNESOTAN

The University Staff Magazine ~ April 1965



## All particle spectrum

Spectrum is the flux as a function of energy.

Flux is the number of particles passing through a given area per unit time.

1 particle
per $\mathrm{km}^{2}$ per millennium

## Minnesota Technolog circa 1966



About 100,000 cosmic ray muons will pass through you each hour.

- Fortunately, muons are "weakly interacting particles" and leave only an ion trail in their wake.
- They do not interact with your nuclei, leaving your basic chemistry unaffected.
- They do cause a background in CCD camera arrays!


## My Professor's plan

- Measure the cosmic ray composition and spectra as a function of latitude at high altitudes on balloons.
- Balloons developed in Minneapolis by people at the physics dept. and a local company known as Winzen.
- Polyethelene resin
- Not affected by uv light
- 0.0001 inches thick
- Small balloons were launched by ourselves from Minneapolis, Devil's Lake, ND, Fayetteville, AK, and in Lake Country in northern MN, north of Hibbing
- Latter payload was lost in a thunderstorm; never found.
- Project Skyhook flights from Ft. Churchill, Manitoba
- Then Tucuman, Argentina \& Queensland, Australia


## Part of team that studied data from the Voyager spacecraft showing it entered the interstellar medium

- 1977 Launch
- 2012 Voyager 1
- 2018 Voyager 2124 AU



## These are not weather balloons



We did use these rubber balloons to find out what the stratospheric winds at above $120,000 \mathrm{ft}$ altitude were doing.

Winds at that altitude reverse direction once a year, setting optimal times for payload launches.

## Sometimes a balloon carries multiple sensors



Upward force from this balloon is less than

## Dissertation hardware The two "goose" experiment



## Join cosmic rays and see the world

- Fort Churchill, Manitoba, Canada
- Skyhook
- Tucuman, Argentina
- Recovery
- Money exchange
- Queensland, Australia
- Townsfolk
- Hydrogen
- Leave taking



## Tucuman, Argentina

- Payload (about 6 ft tall and very thin -1.5 ft diameter) came down in very desert area (no rain for year and half) with thick thorny plants.
- Search plane landed. Crew found young girl who said she would take them to the payload. She crawled through and under leafless thorny bushes. The search crew followed. She found the payload after about 30 minutes. They gave her a $\$ 5$ reward.
- The newspaper from the nearby town had a front page story about how a cash of armaments had been parachuted into the region to supply rebels in the area and how a little girl had been given an enormous bribe to keep it quiet.



## A-frame balloons launched

- We used balloons up to 600,000 cubic ft in volume.
- Balloons went up to about 110,000 ft
- Sometimes the balloon would stall when the temperature changed too rapidly.
- Carried iron ballast to drop, reducing the load and giving it a "push" upward.
- Payloads of up to 1000 lbs of detector, electronics and batteries and ballast (small beads of iron that burn up in the atmosphere).


## Relate story of filling the balloon with an Australian accent in Argentina.







## Gondola 6 ft tall, about $11 / 2 \mathrm{ft}$ in diameter.

Recovery of the payload was often an adventure in itself.



Ed Nye


Later key figure in infrared astronomy


The Wilson cloud chamber picture hints at what I was hired to do at GSFC - something called calorimetry.

## How to measure the energies of very high energy nuclei.



If we can somehow count the number of particles in block of matter, we can get an estimate of the energy.

Relate number of particles to energy by "calibration" at a ground based high energy particle accelerator.

## How to measure the energies of very high energy nuclei.



We used blocks of iron and in between we put layers of scintillators whose light output was proportional to the number of particles.

To identify what nuclei were incident we used scintillators in which the signal was proportional to the square of the charge.

## At NASA we had access to much larger stratospheric balloons



## Requirements

- Large area detectors due to the ever decreasing intensity as energy increased.
- "As long as possible" balloon flights.
- Weight was constrained ( $3000 \mathrm{lbs}-8000 \mathrm{lbs}$ )
- This changed over time as balloon technology improved
- Balloons failed and had to be redesigned
- We had to enclose the apparatus in approximately 1 atmosphere of air to protect electronics.
- We did experiment with apparatus designed to work in zero pressure, but the pressure at which we were expecting to take data was exactly the worst pressure for our light detecting tubes.






Just before payload release. Payload is hanging from Tiny Tim's jaws.

## Source of Balloon Launch videos

## https://www.nasa.gov/scientificballoons/videogallery

Sunrise and Super Tiger videos

